#### Anomalous duality symmetry

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Hermann Nicolai's 60<sup>th</sup> anniversary

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## $\mathcal{N}=8$ supergravity

Is  $\mathcal{N} = 8$  supergravity a consistent quantum field theory?

← Free of ambiguities associated to logarithmic divergences?

Explanation for the excellent ultra-violet behaviour of the 4-graviton amplitudes

- \* Supersymmetry and  $E_{7(7)}$  duality symmetry
- \* Some more hidden symmetry of the quantum theory?

 $\mathcal{N}=8$  supergravity

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[Bern, Carrasco, Dixon, Johansson, Kosower and Roiban]

Maximal supergravity in higher dimensions

$$\sim \int \frac{d^{2\times 7}k}{(2\pi)^{14}} \frac{k^{12}}{k^{14}} \sim p^{12} \ln[\Lambda] \qquad \qquad \mathcal{L}_{7}^{(2)} = e\nabla^{4}R^{4} + \dots$$

$$\int \frac{d^{3\times 6}k}{(2\pi)^{18}} \sim \int \frac{d^{3\times 6}k}{(2\pi)^{18}}$$

$$\int rac{d^{3 imes 6}k}{(2\pi)^{18}} rac{k^{16}}{k^{20}} \sim p^{14} \ln[\Lambda]$$

 $\mathcal{L}_6^{(3)} = e \nabla^6 R^4 + \ldots$ 

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Half-maximal supergravity in higher dimensions

But the explicit computations suggest they are finite. [ Bern, Davies, Dennen, Huang] [ Vanhove, Tourkine]

## Outline

- The SL(2) anomaly
- $R^4$  type invariants in  $\mathcal{N} = 4$  supergravity
- Five dimensions and vector multiplets
- Toward algebraic renormalisation in harmonic superspace

Conclusion

#### [G. Bossard, P. S. Howe and K. S. Stelle, to appear]

- [ G. Bossard, C. Hillmann and H. Nicolai, 1007.5472]
  - G. Bossard, P. S. Howe, K. S. Stelle and P. Vanhove, 1105.6087 ]

## $\mathcal{N}=4$ supergravity

 $\mathcal{N}=4$  supergravity includes

- \* 1 complex scalar field  $\tau$  parametrizing SL(2)/SO(2)
- \* 2 × 4 Dirac fermions  $\chi^i_{\alpha}$
- \* 2 × 6 vectors  $F_{\alpha\beta ij}$
- \* 2 × 4 gravitinos  $\rho_{\alpha\beta\gamma i}$
- \* 2 graviton  $C_{\alpha\beta\gamma\delta}$
- of  $SL(2,\mathbb{C}) \times U(4)$ .

The same letters are used for the corresponding superfields.

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- $\begin{array}{ll} * \ 2 \times 4 \ \text{Dirac fermions } \chi^{i}_{\alpha} & -4 \times 3 \times 1/24 \\ * \ 2 \times 6 \ \text{vectors } F_{\alpha\beta ij} & -6 \times 2 \times 1/6 \\ * \ 2 \times 4 \ \text{gravitinos } \rho_{\alpha\beta\gamma i} & +4 \times 1 \times 7/8 \\ & = 1 \\ * \ 2 \ \text{graviton } C_{\alpha\beta\gamma\delta} & [ \ \text{Marcus} ] \end{array}$

of  $SL(2,\mathbb{C}) \times U(4)$ .

The same letters are used for the corresponding superfields.

 $\mathcal{N}=4$  supergravity's 1-loop anomaly

The rigid SL(2) symmetry is anomalous, and is broken to its parabolic subgroup.

$$\mathbf{f}\,\Gamma^{(1)} = \frac{2+n}{32\pi^2} \int \left(e^{-2\phi}R^{ab}\wedge R_{ab}+\ldots\right)$$

where  $\mathbf{f}\tau = -\tau^2$  with  $\tau = \mathbf{a} + ie^{-2\phi}$ .

Correspondingly at 1-loop

 $\langle g_{\mu
u}(p_1)g_{\sigma
ho}(p_2)a(-p_1-p_2)
angle = rac{2+n}{16\pi^2}arepsilon_{\sigma)(\mu}{}^{\kappa\lambda}(\eta_{
u)(
ho}p_1\cdot p_2-p_{2
u)}p_{1(
ho})p_{1\kappa}p_{2\lambda}$ 

The supersymmetric extension is

$$\mathcal{A}_{\mathbf{f}} = \int \left( e^{-2\phi} R^{ab} \wedge R_{ab} + rac{1}{2} \varepsilon_{abcd} a R^{ab} \wedge R^{cd} + \ldots 
ight)$$

and the Wess-Zumino condition

$$\mathbf{e}\mathcal{A}_{\mathbf{f}} + A_{\mathbf{h}} = 0$$

then implies that

$$\mathcal{A}_{\mathbf{h}} = -\int rac{1}{2}arepsilon_{abcd} R^{ab} \wedge R^{cd}$$

so the rigid symmetry is preserved, but

$$dJ_{\mathbf{h}} = -rac{2+n}{64\pi^2} arepsilon_{abcd} R^{ab} \wedge R^{cd}$$

The supersymmetry invariant  $A_f$  is not defined as a superspace integral, but only as a *d*-closed super 4-form. [Gates]

$$d\mathcal{L}_{4} = \frac{1}{24} E^{E} \wedge E^{D} \wedge E^{C} \wedge E^{B} \wedge E^{A} (D_{A} \mathcal{L}_{BCDE} + 2T_{AB}{}^{F} \mathcal{L}_{FCDE})$$

Focusing on the lowest dimensional components  $([r_1, r_2, r_3]$  such that  $r_1 + 2r_2 + r_3 = 5)$ 

$$\begin{split} D^{p}_{\eta} \mathcal{L}^{ijkl}_{\alpha\beta\gamma\delta} &+ 2 T^{pi}_{\eta\alpha} \,\dot{\varsigma}^{q} \mathcal{L}^{jkl}_{\beta\gamma\delta\varsigma q} + \ \circlearrowright &= 0 \\ D_{\dot{\eta}p} \mathcal{L}^{ijkl}_{\alpha\beta\gamma\delta} &+ 4 D^{i}_{\alpha} \mathcal{L}^{jkl}_{\beta\gamma\delta\dot{\eta}p} + \ \circlearrowright &\approx 0 \\ 2 D_{\dot{\eta}p} \mathcal{L}^{ijk}_{\alpha\beta\gamma\dot{\delta}l} &+ T^{\varsigma}_{\dot{\eta}p\dot{\delta}lq} \mathcal{L}^{ijkq}_{\alpha\beta\gamma\varsigma} + \ \circlearrowright &\approx 0 \\ T_{\dot{\eta}p\dot{\gamma}l}^{\varsigma} \mathcal{L}^{ijq}_{\alpha\beta\varsigma\dot{\delta}l} + \ \circlearrowright &\approx 0 \end{split}$$

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$$d\mathcal{L}_{4} = \frac{1}{24} E^{E} \wedge E^{D} \wedge E^{C} \wedge E^{B} \wedge E^{A} (D_{A} \mathcal{L}_{BCDE} + 2T_{AB}{}^{F} \mathcal{L}_{FCDE})$$

For an arbitrary anti-holomorphic function  $\mathcal{F}[\overline{T}]$ with  $U\overline{U}(1 - T\overline{T}) = 1$ ,  $\begin{pmatrix} U & UT \\ \overline{U}\overline{T} & \overline{U} \end{pmatrix} \in SU(1, 1)$ 

$$\begin{split} \mathcal{L}_{\alpha\beta\gamma\delta}^{ijkl} &= \varepsilon_{\alpha\beta}\varepsilon_{\gamma\delta} \Big( \mathcal{F}[\bar{T}] \mathcal{F}_{\dot{\alpha}\dot{\beta}}^{ij} \mathcal{F}^{\dot{\alpha}\dot{\beta}kl} - \bar{U}^{-2}\bar{\partial}\mathcal{F}[\bar{T}]\varepsilon^{ijpq}\chi_{\dot{\alpha}p}\chi_{\dot{\beta}q} \mathcal{F}^{\dot{\alpha}\dot{\beta}kl} \\ &+ \frac{1}{6}\bar{U}^{-4} \Big(\bar{\partial} - \frac{2T}{1 - T\bar{T}} \Big) \bar{\partial}\mathcal{F}[\bar{T}]\varepsilon^{ijpq}\varepsilon^{klrs}\chi_{\dot{\alpha}p}\chi_{\dot{\beta}q}\chi_{r}^{\dot{\alpha}}\chi_{s}^{\dot{\beta}} \Big) + \circlearrowright \end{split}$$

and

$$\mathcal{L}_{\alpha\beta\gamma\delta l}^{ijk} = -\varepsilon_{\alpha\beta}\varepsilon^{\dot{\eta}\dot{\varsigma}}\chi_{\gamma}^{k}\chi_{\dot{\eta}l} \Big( \mathcal{F}[\bar{\mathcal{T}}]\mathcal{F}_{\dot{\delta}\dot{\varsigma}}^{ij} - \frac{1}{3}\bar{U}^{-2}\bar{\partial}\mathcal{F}[\bar{\mathcal{T}}]\varepsilon^{ijpq}\chi_{\dot{\delta}p}\chi_{\dot{\varsigma}q} \Big) + \bigcirc$$

For  $\mathcal{F}[T] = 1$   $\mathcal{L}_4[1] = \frac{1}{2}R^{ab} \wedge R_{ab} - \frac{i}{4}\varepsilon_{abcd}R^{ab} \wedge R^{cd}$ therefore for  $\mathcal{F}[T] = \tau[T] \equiv i\frac{1-T}{1+T}$  $i\mathcal{L}[\bar{\tau}] - i\bar{\mathcal{L}}[\tau] = \operatorname{Im}[\tau]R^{ab} \wedge R_{ab} + \operatorname{Re}[\tau]\frac{1}{2}\varepsilon_{abcd}R^{ab} \wedge R^{cd} + \dots$ 

defines the anomaly  $\mathcal{A}_{\mathbf{f}}$ .

#### Anomalous Ward identity

At higher order, the anomaly is renormalised consistently

$$\mathbf{f} \, \Gamma_{\epsilon} = rac{1}{16\pi^2} [\mathcal{A} \cdot \Gamma]_{\epsilon}$$

If there would be a logarithmic divergence at 3-loop

$$\mathbf{f}\left(\boldsymbol{\Gamma}^{R}-\frac{\beta_{3}}{\epsilon}\kappa^{4}S^{(3)}\right)\approx\frac{1}{16\pi^{2}}\left(\left[\boldsymbol{\mathcal{A}}\cdot\boldsymbol{\Gamma}\right]^{R}-\frac{\gamma_{2}}{\epsilon}\kappa^{4}\boldsymbol{\mathcal{A}}^{(2)}\right)$$

so that

$$16\pi^2\beta_3\,\mathbf{f}S^{(3)} = \gamma_2\mathcal{A}^{(2)}\;,\qquad \beta_3\,\mathbf{h}S^{(3)} = \mathbf{0}\;,\quad \beta_3\,\mathbf{e}S^{(3)} = \mathbf{0}$$

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$$\left(\frac{\partial}{\partial\mu}-\kappa\frac{\partial}{\partial\kappa}\right)\Gamma\approx-\beta_{3}\kappa^{4}[S^{(3)}\cdot\Gamma]+\mathcal{O}(\kappa^{6})$$

and

$$\left(\frac{\partial}{\partial\mu}-\kappa\frac{\partial}{\partial\kappa}\right)[\mathcal{A}\cdot\mathsf{\Gamma}]\approx-\gamma_{2}\kappa^{4}[\mathcal{A}^{(2)}\cdot\mathsf{\Gamma}]+\mathcal{O}(\kappa^{6})\approx-16\pi^{2}\beta_{3}\kappa^{4}\mathsf{f}[S^{(3)}\cdot\mathsf{\Gamma}]+\mathcal{O}(\kappa^{6})$$

so that

$$16\pi^2eta_3\,{f f}S^{\scriptscriptstyle (3)}=\gamma_2{\cal A}^{\scriptscriptstyle (2)}\;,\qquadeta_3\,{f h}S^{\scriptscriptstyle (3)}=0\;,\quadeta_3\,{f e}S^{\scriptscriptstyle (3)}=0$$

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According to the symmetries, it must be supersymmetric, and invariant with respect to the parabolic subgroup of SL(2).

Therefore we need to study generic invariants of the form

$$\int d^4 x d^{16} \theta \operatorname{Ber}[E] K[T, \overline{T}]$$

or possible generalisations thereof which cannot be written as full-superspace integrals.

#### Normal coordinate expansion

To study the integral

$$\int d^4 x d^{16}\theta \operatorname{Ber}[E] K[T, \bar{T}]$$

it is convenient to integrate over four  $\theta$ 's.

← Requires vector fields in involution.

Only exist in harmonic superspace

 $u'_i \in (U(1) \times U(2) \times U(1)) \setminus U(4)$ 

Then the following vectors are in involution

 $u^{1}{}_{i}\tilde{E}^{i}_{\alpha} , \quad u^{i}{}_{4}\tilde{E}_{\dot{\alpha}i} , \quad D^{1}{}_{r} , \quad D^{r}{}_{4} , \quad D^{1}{}_{4}$ where  $\tilde{E}_{A} = E_{A} - u^{I}{}_{i}\Omega_{A}{}^{i}{}_{j}u^{j}{}_{J}D^{J}{}_{I}$ .

#### Normal coordinate expansion

## $u^{1}{}_{i}\tilde{E}^{i}_{\alpha}$ , $u^{i}{}_{4}\tilde{E}_{\dot{\alpha}i}$ , $D^{1}{}_{r}$ , $D^{r}{}_{4}$ , $D^{1}{}_{4}$

The associated normal coordinate are complex.

 $\mathfrak{u}(4)\cong\overline{\mathbf{1}}^{(-2)}\oplus(2\times\overline{\mathbf{2}})^{(-1)}\oplus\left(\mathfrak{u}(1)\oplus\mathfrak{u}(2)\oplus\mathfrak{u}(1)\right)^{(0)}\oplus(2\times\mathbf{2})^{(1)}\oplus\mathbf{1}^{(2)}$ 

➡ holomorphic normal coordinate expansion

- ★ Usual for Grassmann variables [ Hermann Nicolai, Nucl. Phys. B 140, 294 (1978)]
- \* Not for ordinary commuting variables.

However the normal coordinate expansion factorizes.

- \* The expansion of du depends only on  $z^r_4, z^1_r, z^1_4$
- ★ The expansion of  $Ber[E]K[T, \overline{T}]$  depends only on  $\zeta_1^{\alpha}, \zeta^{\dot{\alpha}4}$

Therefore one can simply expand in the fermionic coordinates only.

#### Normal coordinate expansion

As a result

$$\int d^{4}x d^{16}\theta \operatorname{Ber}[E] \mathcal{K}[T, \overline{T}]$$

$$= \frac{1}{4} \int d\mu_{(4,1,1)} \varepsilon^{\alpha\beta} \varepsilon^{\dot{\alpha}\dot{\beta}} \chi^{1}_{\alpha} \chi^{1}_{\beta} \chi_{\dot{\alpha}4} \chi_{\dot{\beta}4} (\Delta - 2) \Delta \mathcal{K}[T, \overline{T}]$$

$$\sim \int d^{4}x \Big( C^{2} \overline{C}^{2} (\Delta - 2) \Delta \mathcal{K}[\tau, \overline{\tau}] + \dots \Big)$$

It is zero if K is holomorphic, and it is duality invariant for

$$K = -\ln(1 - T\bar{T})$$

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the Kähler potential.

#### Back to the anomaly

All 3-loop candidates can be written as

$$S^{(3)}[G] = \frac{1}{4} \int d\mu_{(4,1,1)} \varepsilon^{\alpha\beta} \varepsilon^{\dot{\alpha}\dot{\beta}} \chi^1_{\alpha} \chi^1_{\beta} \chi_{\dot{\alpha}4} \chi_{\dot{\beta}4} G$$
  
$$\sim \int d^4 x \Big( C^2 \bar{C}^2 G + \dots \Big)$$

So  $\delta S^{\scriptscriptstyle (3)}[G] = S^{\scriptscriptstyle (3)}[\delta G]$  and

$$\mathbf{e}S^{(3)}[G] = \mathbf{h}S^{(3)}[G] = \mathbf{0} \implies G = \mathrm{const}$$

The invariance with respect to the parabolic subgroup implicates duality invariance at this order.

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#### Back to the anomaly

Therefore  $\delta S^{(3)} = 0$  and the only available counter-term satisfying all required symmetries is the duality invariant

$$\int d^4 x d^{16} \theta \operatorname{Ber}[E] \ln(1 - T\bar{T})$$

$$= \frac{1}{2} \int d\mu_{(4,1,1)} \varepsilon^{\alpha\beta} \varepsilon^{\dot{\alpha}\dot{\beta}} \chi^1_{\alpha} \chi^1_{\beta} \chi_{\dot{\alpha}4} \chi_{\dot{\beta}4}$$

$$\sim \int d^4 x \left( C^2 \bar{C}^2 + \ldots \right)$$

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 $\mathcal{N} = 4$  supergravity coupled to *n* vector multiplets

The available invariants are

- ★ 1-loop :
  - \* A duality invariant 1/2 BPS  $F^4$  counter-term
  - \* Infinitely many non-duality invariant pure cocycle in  $R^2$
- ★ 2-loop :
  - \* A 1/4 BPS  $\partial^2 F^4$  counter-term
  - \* A complex non-duality invariant pure cocycle in  $F^2R^2$
- ★ 3-loop :
  - \* A duality invariant 1/4 BPS  $R^4$  counter-term
  - \* A duality invariant full-superspace integral of  $-\ln(1 T \bar{T})$
  - \* Infinitely many non-duality invariant full-superspace integrals.

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 $\mathcal{N}=2$  supergravity in five dimensions

The available invariants are

- ★ 2-loop :
  - \* A duality invariant full-superspace integral of  $\Phi$
  - \* Infinitely many non-duality invariant full-superspace integrals.

$$\int d^{5}x d^{16}\theta \operatorname{Ber}[E] \mathcal{K}[\Phi]$$

$$= \frac{1}{24} \int d\mu_{(4,1)} \varepsilon^{\alpha\beta\gamma\delta} \chi^{1}_{\alpha} \chi^{1}_{\beta} \chi^{1}_{\gamma} \chi^{1}_{\delta} (\partial + 3)(\partial + \frac{3}{2}) \partial (\partial - \frac{3}{2}) \mathcal{K}[\Phi]$$

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## Off-shell formulation of $\mathcal{N} = 4$ supergravity

Counting arguments: can only exist with finitely many auxiliary fields for  $\mathcal{N}=4$  supergravity with 6 vector multiplets [Rivelles and Taylor]

It does exist in ten dimensions for type I supergravity, with  $\tilde{B}_6$  [Howe, Nicolai and Van Proeyen]

In four dimensions SO(6,6) is broken to  $GL(6,\mathbb{R})$  and 15 scalars are realised as 2-form gauge fields.

In four dimensions SO(5,5) is broken to  $GL(5,\mathbb{R})$  and 10 scalars are realised as 3-form gauge fields, 5 vectors as 2-forms.

A duality invariant formulation with n vector multiplets can only exist in harmonic superspace

Off-shell formulation of  $\mathcal{N} = 4$  supergravity

Strong assumption:

It exists a harmonic superspace formulation of  $\mathcal{N} = 4$  supergravity coupled to *n* vector multiplets with all 16 supercharges realised linearly, and for which the action is duality invariant *à la* Henneaux–Teitelboim.

Can we then understand the absence of divergence at 3-loop (respectively 2-loop in five dimensions)?

← algebraic renormalisation in harmonic superspace

## Algebraic renormalisation

The n-loop  $\beta_n$  function associated to an invariant counter-term  $\int \mathcal{L}_{(n)}$  is related to the n-loop anomalous dimension  $\gamma_n$  of the classical Lagrange density  $\mathcal{L}$  as a local operator for mixing with the local operator  $\mathcal{L}_{(n)}$ .

For a first divergence, one proves that

 $(n-1)\beta_n = \gamma_n$ 

#### Algebraic renormalisation

If the Lagrange density is not invariant with respect to a symmetry of the action

$$\delta \mathcal{L} = -D_M \mathcal{L}^M$$
$$\delta \mathcal{L}^M = -D_N \mathcal{L}^{NM}$$
$$\delta \mathcal{L}^{NM} = -D_P \mathcal{L}^{PNM}$$

The sources transforms as forms  $\delta u = 0$ ,  $\delta u_M = -D_M u$ ,  $\delta u_{MN} = -2D_{[M}u_{N]}$ , ... and they all must renormalise consistently

. . .

. . .

$$\mathcal{L}_{\flat} = \mathcal{L} + \gamma_n \ln(\frac{\Lambda}{\mu}) \mathcal{L}_{\scriptscriptstyle (n)} + \dots$$

$$\mathcal{L}_{\flat}^M = \mathcal{L}^M + \gamma_n \ln(\frac{\Lambda}{\mu}) \mathcal{L}_{\scriptscriptstyle (n)}^M + \dots$$

$$\mathcal{L}_{\flat}^{NM} = \mathcal{L}^{NM} + \gamma_n \ln(\frac{\Lambda}{\mu}) \mathcal{L}_{\scriptscriptstyle (n)}^{NM} + \dots$$

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## $\mathcal{N}=4$ counter-term density

We have seen that the only available duality invariant full-superspace integral at 3-loop is the integral of the Kähler potential.

Under duality

$$\delta \Big( \mathsf{Ber}[E](-1)\mathsf{ln}\big( \frac{-i}{2}(\tau - \overline{\tau}) \big) \Big) = -2h\mathsf{Ber}[E] + f\mathsf{Ber}[E](\tau + \overline{\tau})$$

Within an off-shell formulation,

 $Ber[E] = D_M \Psi^M[E, V]$ 

and respectively for  $Ber[E](\tau + \overline{\tau})$ , where V is some prepotential for the supergravity fields.

→ Whereas explicit prepotential are forbidden within the background field method.

#### Conclusion

Can supersymmetry and duality invariance explain the absence of logarithm divergence at 3-loop in  $\mathcal{N} = 4$  d = 4 supergravity, respectively 2-loop in 5 dimensions?

With the rather strong assumption that an off-shell harmonic formulation with a duality invariant action exists

If yes, what about  $\mathcal{N}=8$  supergravity?

$$\int \partial^8 R^4 + \cdots \sim \int d^4 x d^{28} \theta \mathcal{E} \chi^4 \stackrel{\sim}{\sim} \int d^4 x d^{32} \theta \operatorname{Ber}[E] \mathcal{K}[\phi]$$

But asking for an off-shell formulation with all 32 supercharges does not seem reasonable then.

# Happy Birthday Hermann !!