

# The Gravity/Coset Conjecture and a Possible Algebraic Description of Emergent Space

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work with  
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Nicolai Fest, 6–8 September, 2012  
AEI, Golm

# Gravity/Coset Correspondence

$E_{10}$  : Damour, Henneaux, Nicolai '02;  
related: Ganor '99 '04;  $E_{11}$  : West '01

SUGRA<sub>11</sub> (OR M-THEORY)

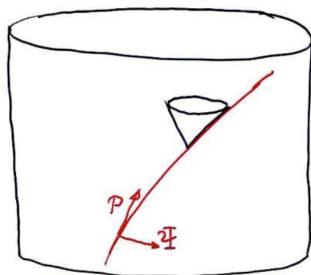
$$G_{\mu\nu}(t, \vec{x})$$

$$A_{\mu\nu\lambda}(t, \vec{x})$$

$$\psi_{\mu}(t, \vec{x})$$



MASSLESS SPINNING PARTICLE  
ON COSET  $E_{10}/K(E_{10})$



# Gravity/Coset Correspondence

$$G_{\mu\nu}(t, \mathbf{x}), \mathcal{A}_{\mu\nu\lambda}(t, \mathbf{x}), \psi_\mu(t, \mathbf{x})$$

$$g(t) = \exp(h_b^a(t) K_a^b) \exp\left[\frac{1}{3!} A_{abc}(t) E^{abc} + \frac{1}{6!} A_{a_1 \dots a_6}(t) E^{a_1 \dots a_6} + \frac{1}{9!} A_{a_0 | a_1 \dots a_8}(t) E^{a_0 | a_1 \dots a_8} + \dots\right]$$

$$S_{11} = \int d^4x \left\{ \frac{E}{4} R(G) - \frac{E}{48} (d\mathcal{A}_3)^2 + \dots \right\}$$

$$S_1^{\text{COSET}} = \int dt \left\{ \frac{1}{4n(t)} \langle P(t), P(t) \rangle - \frac{i}{2} (\Psi(t) | \mathcal{D}^{\text{vs}} \Psi(t))_{\text{vs}} + \dots \right\}$$

Gradient Expansion (BKL)  
(~ Small Tension Expansion:  
 $\alpha' \rightarrow \infty$ )

Height Expansion  
in Kac-Moody Algebra

$$\partial_{x^1}^{k_1} \partial_{x^2}^{k_2} \dots \partial_{x^{10}}^{k_{10}} \ll \partial_T^{k_1+k_2+\dots+k_{10}}$$

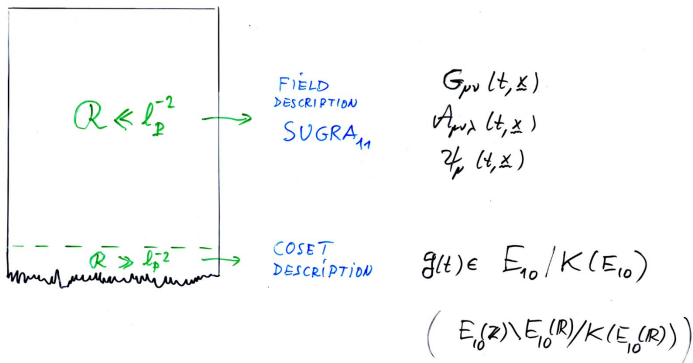
Root:

$$\alpha = n_1 \alpha_1 + n_2 \alpha_2 + \dots + n_{10} \alpha_{10}$$

# Idea: Two 'dual' or 'complementary' descriptions

GRAVITY / COSET CORRESPONDENCE

NEAR SPACELIKE SINGULARITY



The 'singularity' is 'resolved' by the effective 'disappearance' of space, and the replacement of dynamical fields,  $g_{ij}(t, \mathbf{x}), A_{ijk}(t, \mathbf{x}), \dots$  by a **Lie-algebraic** variable  $g(t) \in E_{10}/K_{10}$

# KAC-MOODY ALGEBRAS

$SU(2)$   
 $\cong$   
 $A_1$   
 $\cong$   
 $SL(2)$

$$[J_z, J_+] = +J_+ \quad [J_z, J_-] = -J_-$$

↑ CARTAN GENERATOR (DIAGONAL)      ↑ RAISING GENERATOR      ↑ CARTAN GENERATOR      ↑ LOWERING GENERATOR

CARTAN SUBALGEBRA : LINEAR SPACE  $\mathbb{R}^r$        $\overset{\text{RANK}}{\uparrow} r$

$$\mathfrak{h} = \{ \beta^a h_a ; a=1, \dots, r \}$$

$[h', h''] = 0$

$\uparrow r$  independent Cartan generators  
 coordinates in Cartan space:  $h = \sum_a \beta^a h_a$

TRIANGULAR DECOMPOSITION:

$$\mathfrak{g} = \mathfrak{n}^- \oplus \mathfrak{h} \oplus \mathfrak{n}^+$$

↑ LOWERING GENERATORS  $F_\alpha$       ↑ CARTAN  $h$       ↑ RAISING GENERATORS  $E_\alpha$

$$[h, E_\alpha^{(s)}] = \alpha(h) E_\alpha^{(s)}$$

Cartan  
 $h = \sum_a \beta^a h_a$

↑  
 RAISING  
 GENERATOR(S)

↑  
 ROOT  $\equiv$

EIGENVALUE OF  $\text{ad}_h$

AS A LINEAR FORM OF  $h \in \text{CSA}$

$$h = \beta^a h_a \rightarrow \alpha(h) = \alpha_a \beta^a \equiv \alpha(\beta)$$

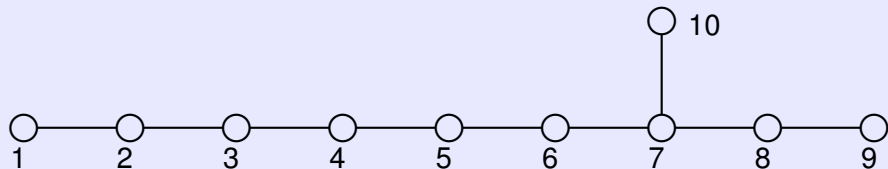
$$[E_\alpha^{(s)}, E_\beta^{(t)}] = C_{\alpha\beta}^{(s+t)} E_{\alpha+\beta}^{(s+t)}$$

$$[h, F_\alpha^{(s)}] = -\alpha(h) F_\alpha^{(s)}$$

↑  
 LOWERING  
 GENERATORS :  $F_\alpha^{(s)} \equiv E_{-\alpha}^{(s)}$

+ JACOBI + SERRE RELATIONS

# $E_{10}$ Dynkin Diagram (= Cartan Matrix)

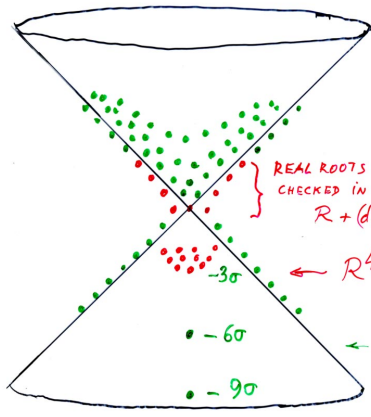


**Figure:** *Dynkin diagram of  $E_{10}$  with numbering of nodes.*

Here, leftmost  $E_{10}$  “hyperbolic” node labelled 1

# $E_{10}$ Root Diagram

ROOTS OF  $E_{10}$





# Kinematics on Coset $G/K$

$$\mathcal{V}(t) \in G/K$$

$$\mathbf{v} \equiv \partial_t \mathcal{V} \mathcal{V}^{-1} \in \text{Lie}(G)$$

decomposed into  $\mathbf{v} = \mathcal{P} + \mathcal{Q}$  where

$$\mathcal{P} = \mathbf{v}^{\text{sym}} = \frac{1}{2}(\mathbf{v} + \mathbf{v}^T) : \text{'coset velocity'}$$

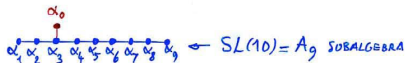
$$\mathcal{Q} = \mathbf{v}^{\text{antisym}} = \frac{1}{2}(\mathbf{v} - \mathbf{v}^T) : \text{'K angular velocity'}$$

Coset Action :

$$S_{\text{BOS}}^{\text{coset}} = \int \frac{dt}{n(t)} \frac{1}{4} \langle \mathcal{P}(t), \mathcal{P}(t) \rangle$$


$n(t)$  : coset lapse  $\rightarrow$  constraint  $\langle \mathcal{P}(t), \mathcal{P}(t) \rangle = 0$


# DECOMPOSING $E_{10}$ WRT. $GL(10)$ SUBALGEBRA 4




"LEVEL"  $l$ :  $\alpha = l \alpha_0 + \sum_{j=1}^9 m_j \alpha_j$

$l=0$   $GL(10)$  GENERATORS  $K^a_b$   $[K^a_b, K^c_d] = \delta^c_b K^a_d - \delta^a_d K^c_b$

$l = \pm 1$   $E^{[a_1 a_2 a_3]}$ ,  $F_{[a_1 a_2 a_3]}$   3 INDICES

$l = \pm 2$   $E^{[a_1 \dots a_6]}$ ,  $F_{[a_1 \dots a_6]}$   6 INDICES

$l = \pm 3$   $E^{[a_0 a_1 \dots a_8]}$ ,  $F_{[a_0 a_1 \dots a_8]}$   9 INDICES

$l = \pm 4$    $\oplus$   12 INDICES

# EXPLICIT PARAMETRIZATION OF $E_{10}/K(E_{10})$ <sup>8</sup>

$$g(t) = e^{h_a^b(t) K_a^b} e^{\frac{1}{2!} A_{q_1 q_2 a_3} E^{q_1 a_2 a_3} + \frac{1}{6!} A_{a_1 \dots a_6} E^{a_1 \dots a_6} + \frac{1}{9!} A_{a_1 a_2 a_3 a_4 a_5 a_6} E^{a_1 a_2 a_3 a_4 a_5 a_6} + \dots}$$

$\uparrow$   $GL(10): K_a^b$   
 $\downarrow h_a^b$   
 $g^{ab}(t) = (e^h)^a_c (e^h)^b_c$

$A_{q_1 q_2 a_3}$   $A_{a_1 \dots a_6}$   $A_{a_1 a_2 a_3 a_4 a_5 a_6}$  + ...

indices raised by  $g^{ab}$

$$\int_1^{E_{10}/K(E_{10})} = \int \frac{dt}{m(t)} \left[ \frac{1}{4} (g^{ac} g^{bd} - g^{ab} g^{cd}) \dot{g}_{ab} \dot{g}_{cd} + \frac{1}{2 \cdot 3!} \dot{A}_{q_1 q_2 a_3} \dot{A}^{q_1 a_2 a_3} + \frac{1}{2 \cdot 6!} \dot{D} A_{a_1 \dots a_6} \dot{D} A^{a_1 \dots a_6} + \frac{1}{2 \cdot 9!} \dot{D} A_{a_1 a_2 a_3 a_4 a_5 a_6} \dot{D} A^{a_1 a_2 a_3 a_4 a_5 a_6} + \dots \right]$$

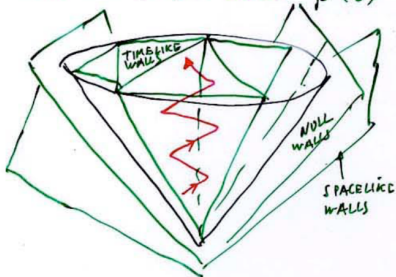
$$D A_{q_1 \dots a_6} = \dot{A}_{q_1 \dots a_6} + 10 A_{[q_1 \dots q_3] a_4 \dots a_6} \dot{A}_{q_1 \dots a_6}$$

$$D A_{a_1 a_2 a_3 a_4 a_5 a_6} = \dot{A}_{a_1 a_2 a_3 a_4 a_5 a_6} + 42 A_{\langle a_1 \dots a_3 \rangle a_4 \dots a_6} \dot{A}_{a_1 \dots a_6} - 42 \dot{A}_{\langle a_1 \dots a_3 \rangle a_4 \dots a_6} A_{a_1 \dots a_6} + 280 A_{\langle a_1 \dots a_3 \rangle a_4 \dots a_6} \dot{A}_{a_1 \dots a_6}$$

# Correspondence Between Coset/Gravity Dynamics

$$\mathcal{L}_{10} \sim (\dot{g}^{-1} \dot{g})^2 + (\dot{A}_3)^2 + (\dot{A}_6 + A_3 \dot{A}_3)^2 + (\dot{A}_9 + A_6 \dot{A}_3 + A_3 A_3 \dot{A}_2)^2 + \dots$$

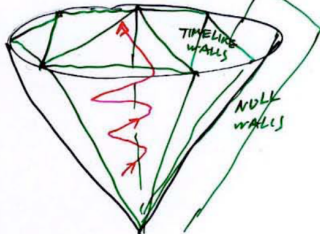
BILLIARD WITH INFINITE NUMBER OF EXPONENTIAL WALLS FOR CARTAN ELEMENT  $\beta^{\dot{A}}(t)$



$$\mathcal{L}_{11}^{\text{SUGRA}} = \int d^2x \left[ \sqrt{-G} \left( R(G) - \frac{(dA_3)^2}{48} \right) + \frac{1}{(12)^4} F_4 \wedge F_4 \wedge A_3 \right]$$

$F_4 = d \wedge dt_3$

BILLIARD WITH LARGE BUT FINITE # OF EXPONENTIAL WALLS FOR  $\beta^{\dot{A}}(t, x)$ , DIAGONAL PART OF  $G_{ij}(t, x)$  IN IWASAWA DECOMP.



# Gravity/Coset Dictionary

$$\mathcal{H}_1 = \frac{1}{2} G^{\mu\nu} \pi_\mu \pi_\nu + \sum_{\alpha} c_{\alpha}(\rho, p) e^{-2\alpha(\beta)}$$

$$\alpha(\beta) = \sum_i m_i \alpha_i(\beta)$$

↑  
 $m_i \in \mathbb{N}$  SIMPLE ROOTS

$$\mathcal{H}_0 = \frac{1}{2} G^{\mu\nu} \pi_\mu \pi_\nu + \sum_A c_A(\rho, p, \beta_1, \beta_2, \dots) e^{-2w_A(\beta)}$$

$$w_A(\beta) = \sum_i m_i w_i(\beta)$$

↑  
DOMINANT WALLS

DICTIONARY

$$g^{ab}(t) = (e^h)_c^a (e^h)_c^b = G^{ab}(t, \vec{x}_0)$$

WRT A SPECIAL FRAME

$$A_{q_1 q_2 q_3}(t) = F_{0 q_1 q_2 q_3}(t, \vec{x}_0)$$

$$\theta^q(x) = e^q(x) dx^i$$

$$DA^{q_1 \dots q_6}(t) = g^{q_1 a_1} g^{q_2 a_2} \dots g^{q_6 a_6} [A_{q_1 \dots a_6} + 10 A_{[3} A_{3]}] = -\frac{1}{4!} \varepsilon^{q_1 \dots q_6 b_1 \dots b_4} F_{b_1 \dots b_4}(t, \vec{x}_0)$$

$$DA^{b_1 \dots b_4}(t) = g^{b_1 a_1} \dots g^{b_4 a_4} [A_{a_1 \dots a_4} + 42 A_3 A_3 + 250 A_3 A_3 A_3] = +\frac{3}{2} \varepsilon^{q_1 \dots q_6 b_1 b_2} C^b(\vec{x}_0)$$

$$\uparrow \quad \uparrow \quad \uparrow$$

$$d\theta^q = \frac{1}{2} C^q_{bc} \theta^b \theta^c$$

THE CORRESPONDENCE WORKS FOR ALL TERMS OF HEIGHT  $\leq 29$

$$\sum_i m_i \leq 29$$

$$\sum_i m_i \leq 29$$

# Classical Coset Hamiltonian

Bosonic

$$S_1 = \int \frac{dt}{2n(t)} \|(\partial_t \mathcal{V} \mathcal{V}^{-1})^{\text{Sym}}\|^2$$

Iwasawa:

$$\mathcal{V}(t) = \exp(\underbrace{\beta^a(t) H_a}_{\text{Cartan}}) \exp\left(\underbrace{\sum_{\alpha>0} v^\alpha(t) E_\alpha}_{\text{Raising (+ multiplicity)}}\right)$$

$$S^{\text{coset}} = \int \frac{dt}{n(t)} \left[ \frac{1}{2} G_{ab} \dot{\beta}^a \dot{\beta}^b + \sum_{\alpha>0} \frac{1}{4} e^{2\alpha(\beta)} (\dot{v}^\alpha + c v^{\alpha'} \dot{v}^{\alpha''} + \dots)^2 \right]$$

$$H^{\text{coset}}(\beta^a, \pi_a; v^\alpha, p_\alpha) = n(t) \left[ \frac{1}{2} G^{ab} \pi_a \pi_b + \sum_{\alpha>0} e^{-2\alpha(\beta)} \Pi_\alpha^2(p, v) \right]$$

# Quantum Gravity $\leftrightarrow$ Quantum Coset Model

Quantum Coset Model: in configuration space  $\beta^a, \nu^\alpha$

$$\square_{E_{10}/K_{10}} \Psi(\beta^a, \nu^\alpha) = 0$$

$$\left[ -G^{ab} \partial_{\beta^a} \partial_{\beta^b} - \sum_{\alpha > 0} e^{-2\alpha(\beta)} \partial_{\nu^\alpha}^2 + \dots \right] \Psi(\beta^a, \nu^\alpha) = 0$$

Infinite-dimensional Klein-Gordon type equation:  $- + + + + \dots$

From Hull-Townsend '95 expect  $E_{10}(\mathbb{Z})$  symmetry, i.e.

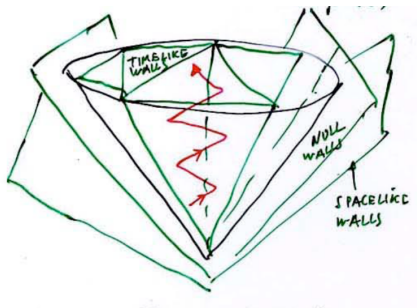
$\Psi(\beta, \nu) =$  automorphic function over  $E_{10}(\mathbb{Z}) \backslash E_{10}(\mathbb{R}) / K_{10}(\mathbb{R})$

(Ganor '99; Brown, Ganor, Helfgott '04)

# Baby Version of the Coset Model: Billiard

Particle in  $\beta$  (Cartan) space; billiard walls = simple roots of  $E_{10}$   
(Damour-Henneaux '00)

coset billiard = BKL gravity billiard  $\simeq$  Kasner mini-superspace  
 $g_{ab}(t) = e^{-2\beta^a(t)} \delta_{ab} + \text{effect of leading spatial gradients}$





# Baby Version of the Quantum Coset Model: Quantum Billiard

Kleinschmidt, Nicolai et al. '09 '10

Setting  $n = 1$  one has to quantize

$$\mathcal{L} = \frac{1}{2} \sum_{a,b=1}^d \dot{\beta}^a G_{ab} \dot{\beta}^b = \frac{1}{2} \left[ \sum_{a=1}^d (\dot{\beta}^a)^2 - \left( \sum_{a=1}^d \dot{\beta}^a \right)^2 \right]$$

with null constraint  $\dot{\beta}^a G_{ab} \dot{\beta}^b = 0$  on billiard domain.

Canonical momenta:  $\pi_a = G_{ab} \dot{\beta}^b \Rightarrow \mathcal{H} = \frac{1}{2} \pi_a G^{ab} \pi_b.$

**Wheeler–DeWitt** (WDW) equation in canonical quantization

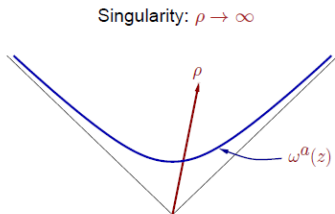
$$\mathcal{H}\Psi(\beta) = -\frac{1}{2} G^{ab} \partial_a \partial_b \Psi(\beta) = 0$$

Klein–Gordon ‘inner product’.

# Quantum Billiard: Polar Coordinates $\beta^a = \rho \omega^a(z)$

Introduce new coordinates  $\rho$  and  $\omega^a(z)$  from 'radius' and coordinates  $z$  on unit hyperboloid

$$\beta^a = \rho \omega^a, \quad \omega^a G_{ab} \omega^b = -1$$
$$\rho^2 = -\beta^a G_{ab} \beta^b$$



Timeless WDW equation in these variables

$$\left[ -\rho^{1-d} \frac{\partial}{\partial \rho} \left( \rho^{d-1} \frac{\partial}{\partial \rho} \right) + \rho^{-2} \Delta_{\text{LB}} \right] \Psi(\rho, z) = 0$$

Laplace–Beltrami operator on unit hyperboloid

# Solving the WDW Equation

$$\left[ -\rho^{1-d} \frac{\partial}{\partial \rho} \left( \rho^{d-1} \frac{\partial}{\partial \rho} \right) + \rho^{-2} \Delta_{\text{LB}} \right] \Psi(\rho, z) = 0$$

Separation of variables:  $\Psi(\rho, z) = R(\rho)F(z)$

For

$$-\Delta_{\text{LB}} F(z) = E F(z)$$

get

$$R_{\pm}(\rho) = \rho^{-\frac{d-2}{2} \pm i \sqrt{E - \left(\frac{d-2}{2}\right)^2}}$$

[Positive frequency coming out of singularity is  $R_-(\rho)$ .]

Left with spectral problem on hyperbolic space.

# Boundary Conditions

Classical Billiard suggests Dirichlet boundary conditions at time-like walls

Then yields a Weyl ( $E_{10}$ )-modular form by using Weyl images

Interesting *arithmetic* aspects of this Weyl billiard

$$W^+(E_{10}) = \text{PSL}_2(O) \quad (\text{Feingold, Kleinschmidt, Nicolai '08})$$

even Weyl group                      octavians = integer octonions

generalizes  $W^+(AE_3) = \text{PSL}_2(\mathbb{Z})$  for 3 + 1 gravity

$$\longrightarrow \Psi_{E_{10}}^{\text{billiard}} = \text{Maass wave form for } \text{PSL}_2(O)$$

Deeper issue of space like boundary condition at  $\rho \rightarrow \infty$  (singularity)

? need some final-state boundary condition ?

(Horowitz, Maldacena '04)

# (Bosonic) Evolution Equations and Constraints

Supergravity	Coset	Name
$\mathcal{G}_{ab} = 0$	$\mathcal{D}P_{ab}^{(0)} = 0$	Einstein dynamical eq.
$\mathcal{M}^{a_1 a_2 a_3} = 0$	$\mathcal{D}P_{a_1 a_2 a_3}^{(1)} = 0$	Matter dynamical eq.
$D_{[0} F_{a_1 \dots a_4]} = 0$	$\epsilon_{a_1 \dots a_4 b_1 \dots b_6} \mathcal{D}P_{b_1 \dots b_6}^{(2)} = 0$	F-Bianchi I
$R_{[0 a b] c} = 0$	$\epsilon_{b c d_1 \dots d_8} \mathcal{D}P_{a   d_1 \dots d_8}^{(3)} = 0$	R-Bianchi I
$\mathcal{G}_{00} = 0$	$\langle \mathcal{P}   \mathcal{P} \rangle = 0$	Hamiltonian constraint
$\mathcal{G}_{0a} = 0$	$\epsilon_{a c_1 \dots c_9} \mathcal{C}_{c_1 \dots c_9}^{(3)} = 0$	Momentum constraint
$\mathcal{M}^{0 a_1 a_2} = 0$	$\epsilon_{b_1 \dots b_{10}} \mathcal{C}_{b_1 \dots b_{10}    a_1 a_2}^{(4)} = 0$	Gauss constraint
$D_{[c_1} F_{c_2 \dots c_5]} = 0$	$\epsilon_{b_1 \dots b_{10}} \epsilon_{a_1 \dots a_5 c_1 \dots c_5} \mathcal{C}_{b_1 \dots b_{10}    a_1 \dots a_5}^{(5)} = 0$	F-Bianchi II
$R_{[c_1 c_2 c_3] a_0} = 0$	$\epsilon_{b_1 \dots b_{10}} \epsilon_{a_1 \dots a_7 c_1 c_2 c_3} \mathcal{C}_{b_1 \dots b_{10}    a_0   a_1 \dots a_7}^{(6)} = 0$	R-Bianchi II

# Structure of Known (Bosonic) Constraints

Hamiltonian constraint:  $\mathcal{G}_{00} = 0 \rightarrow \langle \mathcal{P} | \mathcal{P} \rangle = 0$

Other SUGRA constraints in  $GL(10)$  decomposition

Momentum: 
$$\mathcal{C}_{a_1 \dots a_9}^{(3)} = P_{ca_1}^{(0)} P_{c|a_2 \dots a_9}^{(3)} + 28 P_{a_1 a_2 a_3}^{(1)} P_{a_4 \dots a_9}^{(2)},$$

Gauss: 
$$\mathcal{C}_{b_1 \dots b_{10} || a_1 a_2}^{(4)} = P_{a_1 b_1 b_2}^{(1)} P_{a_2 | b_3 \dots b_{10}}^{(3)} + \frac{21}{5} P_{a_1 b_1 \dots b_5}^{(2)} P_{a_2 b_6 \dots b_{10}}^{(2)},$$

Bianchi-F: 
$$\mathcal{C}_{b_1 \dots b_{10} || a_1 \dots a_5}^{(5)} = P_{a_1 \dots a_4 b_1 b_2}^{(2)} P_{a_5 | b_3 \dots b_{10}}^{(3)},$$

Bianchi-R: 
$$\mathcal{C}_{b_1 \dots b_{10} || a_0 | a_1 \dots a_7}^{(6)} = P_{a_0 | b_1 \dots b_8}^{(3)} P_{b_9 | b_{10} a_1 \dots a_7}^{(3)}.$$

# The $E_{10}$ Noether Current

Global  $E_{10}$  symmetry of coset action  $\Rightarrow$  existence of infinite number of Noether charges:  $\mathcal{J} = n^{-1} \mathcal{V}^{-1} \mathcal{P} \mathcal{V}$ .

Expansion of  $\mathcal{J} \in \text{Lie}(E_{10})$  along generators (with truncation  $\mathcal{J}^{(\ell)} = 0$  for  $\ell = -4, -5, -6, \dots$ )

$$\begin{aligned} \mathcal{J} = & \frac{1}{9!} \mathcal{J}^{(-3)m_0|m_1\dots m_8} F_{m_0|m_1\dots m_8} + \frac{1}{6!} \mathcal{J}^{(-2)m_1\dots m_6} F_{m_1\dots m_6} + \frac{1}{3!} \mathcal{J}^{(-1)mnp} F_{mnp} + \\ & + \mathcal{J}^{(0)n} K^m_n + \frac{1}{3!} \mathcal{J}^{(1)mnp} E^{mnp} + \frac{1}{6!} \mathcal{J}^{(2)m_1\dots m_6} E^{m_1\dots m_6} + \dots \end{aligned}$$

$$\mathcal{J}^{(-3)m_0|m_1\dots m_8}$$

$$= P^{(3)m_0|m_1\dots m_8},$$

$$\mathcal{J}^{(-2)m_1\dots m_6}$$

$$= P^{(2)m_1\dots m_6} + \frac{1}{3!} A_{pqr} P^{(3)p|qrm_1\dots m_6},$$

$$\mathcal{J}^{(-1)mnp}$$

$$\begin{aligned} = & P^{(1)mnp} + \frac{1}{3!} A_{rst} P^{(2)rstmnp} + \\ & + \left( \frac{2}{3} A_{r_1\dots r_6} + \frac{1}{72} A_{r_1 r_2 r_3} A_{r_4 r_5 r_6} \right) P^{(3)r_1|r_2\dots r_6 mnp}. \end{aligned}$$

# Hidden Sugawara Structure of Constraints: J J

Remarkably the constraints admit a formulation in terms of conserved Noether charges  $\mathcal{J}$ : Hamiltonian constraint:  $\mathcal{L} \equiv \langle \mathcal{J} | \mathcal{J} \rangle \approx 0$  ( $E_{10}$  singlet), and :

$$\begin{aligned}
 \text{Momentum:} \quad \mathcal{L}^{(-3)^{n_1 \dots n_9}} &= 28 \mathcal{J}^{(-1)^{n_1 n_2 n_3} (-2)^{n_4 \dots n_9}} \\
 &+ \mathcal{J}^{(0)^{n_1} (-3)^{p|n_2 \dots n_9}} \\
 \text{Gauss:} \quad \mathcal{L}^{(-4)^{m_1 \dots m_{10}|n_1 n_2}} &= \frac{21}{5} \mathcal{J}^{(-2)^{n_1 m_1 \dots m_5} (-2)^{n_2 m_6 \dots m_{10}}} \\
 &+ \mathcal{J}^{(-1)^{n_1 m_1 m_2} (-3)^{n_2 | m_3 \dots m_{10}}}, \\
 \text{Bianchi-F:} \quad \mathcal{L}^{(-5)^{m_1 \dots m_{10}|n_1 \dots n_5}} &= \mathcal{J}^{(-2)^{n_1 \dots n_4 m_1 m_2} (-3)^{n_5 | m_3 \dots m_{10}}} \\
 \text{Bianchi-R:} \quad \mathcal{L}^{(-6)^{m_1 \dots m_{10}|n_0 | n_1 \dots n_7}} &= \mathcal{J}^{(-3)^{n_0 | m_1 \dots m_8} (-3)^{m_9 | m_{10} n_1 \dots n_7}}
 \end{aligned}$$



# Analogy: String / Kac-Moody Coset

Gauge-fixed action	$S = \int \int d\tau d\sigma (\dot{X}^2 - X'^2)$	$S = \int \frac{dt}{n(t)} [(\dot{g}g^{-1})^{\text{sym}}]^2$
Infinite symmetry	$\delta X^\mu = \epsilon_n^\mu e^{in(\tau \mp \sigma)}$	$g(t) \rightarrow k(t) g(t) g_0$
Infinite # conserved Noether charges	$j_n^\mu$	$\mathcal{J} = g^{-1} v^{\text{sym}} g / n$
Integrable	$X_L^\mu = \frac{i\ell}{2} \sum \frac{j_n^\mu}{n} e^{-in(\tau - \sigma)}$	$g(t) = e^{t\mathcal{J}} g(0) k(t)$
$\infty$ # constraints	$L_m \sim \int d\sigma e^{-im\sigma} (\partial_- X)^2$	SUGRA constraints inc. spatial gradients
Sugawara structure	$L_m = \frac{1}{2} \sum j_{m-n}^\mu j_n^\mu$	YES, DKN'07'09
Affine KM alg.	$[j_m^a, j_n^b] = f_c^{ab} j_{m+n}^c + O(K)$	Hyperbolic KM alg.
Constraint alg.	$[L_m, L_n] = (m - n)L_{m+n} + O(c)$	$\{J_\alpha, J_\beta\} = f_{\alpha\beta}^{\alpha+\beta} J_{\alpha+\beta}$ Hyperbolic analog of Virasoro ?

# Towards a Generalized Virasoro Algebra

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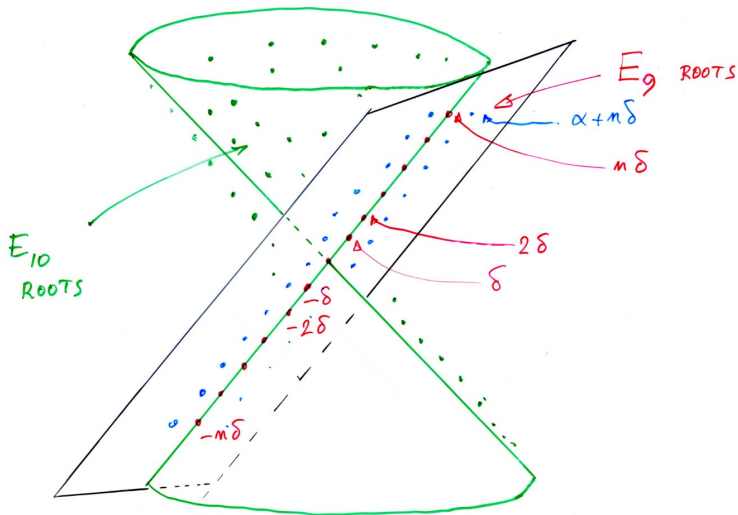
- Usual Virasoro constraints  $L_n$  can be considered as labelled by a string of null roots  $L_{n\delta} := \sum_{\beta \in \Delta^{\text{aff}}} T_{n\delta - \beta} T_{\beta}$  in terms of affine  $K^M$  generators  $T_{\alpha}$ :  $[T_{\alpha_1}, T_{\alpha_2}] = f_{\alpha_1 \alpha_2}^{\alpha_1 + \alpha_2} T_{\alpha_1 + \alpha_2} + \kappa_{\alpha_1, \alpha_2} K$

- Similarly one can label the Gravity/Coset constraints  $\mathcal{L}$  in terms of  $E_{10}$  roots.

e.g.  $GL(10)$  multiplet  $\mathcal{L}^{(-3)^{[n_1 \dots n_9]}}$ : highest weight  $\mathcal{L}^{(-3)^{2345678910}}$ ,  
 associated with **negative, null** root (which is the fundamental weight  $\Lambda_1$  associated with the leftmost  $E_{10}$  “hyperbolic” node)

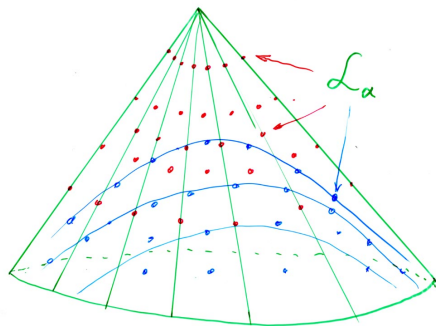
$$\alpha = \Lambda_1 = -\delta^{(3)} = -(\alpha_2 + 2\alpha_3 + 3\alpha_4 + 4\alpha_5 + 5\alpha_6 + 6\alpha_7 + 4\alpha_8 + 2\alpha_9 + 3\alpha_{10})$$

# Affine Kac-Moody ( $E_9$ ) versus Hyperbolic ( $E_{10}$ )



# Full Root Support of Coset Constraints

Expect: root support of fundamental rep.  $L(\Lambda_1)$  + spatial gradients which add multiples  $-n\delta$



root support = cone of all past-directed null and time-like roots

# 'Sugawara Content' of Constraints

$$\mathcal{L}_\alpha = \sum_{\beta, s, s'} M_{s, s'}(\alpha, \beta) \mathcal{J}_{\alpha - \beta}^{(s)} \mathcal{J}_\beta^{(s')}$$

? numerical coefficients:  $M_{s, s'}(\alpha, \beta)$

? 'constituent roots':  $\alpha = \beta_1 + \beta_2$  with  $\beta_1 = \alpha - \beta$ ,  $\beta_2 = \beta$

# Universality vs Non-Universality of Constraints

While  $[\mathcal{L}_{IIA}]_{\text{real roots}} = [\mathcal{L}_{IIB}]_{\text{real roots}}$ ,  $[\mathcal{L}_{IIA}]_{\text{imaginary roots}} \neq [\mathcal{L}_{IIB}]_{\text{imaginary roots}}$   
because of different manifest **covariantization groups**

II A



EXPLICITLY COVARIANT UNDER  
 $A_9 = SL(10)$

II B



EXPLICITLY COVARIANT UNDER  
 $A_8 \oplus A_1 : SL(9) \times SL(2)$

## Universality vs Non-Universality of Constraints (2)

First possibility: try to combine the different covariantizations to define 'more complete' coset constraints: 'see saw' technique.

Second possibility: admit that the two different choices of (maximal) **parabolic** subgroups

$$IIA = \text{Lie}(E_{10}) - \{f_{10}\} \neq IIB = \text{Lie}(E_{10}) - \{f_8\}$$

correspond to **two different 'gauge fixings'** of the underlying  $E_{10}$ -invariant coset action  $\Rightarrow$  two different additional constraints

$$S_{IIA} = \int dt \left[ \frac{\|P\|^2}{n(t)} + \lambda_{IIA}^\alpha(t) \mathcal{L}_\alpha^{IIA} \right] \neq S_{IIB} = \int dt \left[ \frac{\|P\|^2}{n(t)} + \lambda_{IIB}^\alpha(t) \mathcal{L}_\alpha^{IIB} \right]$$

# Algebra of constraints: Generalized Virasoro alg.

For each choice of (maximal) **parabolic** subgroup (IIA or IIB)  $\leftrightarrow$  associated **level** decomposition ( $\ell_{IIA}$  counts  $e_{10}$ ,  $\ell_{IIB}$  counts  $e_8$ ), and

$$\begin{aligned} \mathcal{L}^{-(\ell+1)} &= \{ \mathcal{J}^{(-1)}, \mathcal{L}^{-(\ell)} \} \\ \mathcal{L}^{-(\ell)} &= \sum_{p+q=\ell} \mathcal{J}^{(-p)} \cdot \mathcal{J}^{(-q)} \end{aligned}$$

$\Rightarrow$  **Hyperbolic** generalization of usual **Virasoro**, and (hopefully) of gravity's **diffeomorphism** algebra:

$$\{ \mathcal{L}_\alpha, \mathcal{L}_\beta \} \sim \sum_\gamma \mathcal{J}_{\alpha+\beta-\gamma} \mathcal{L}_\gamma$$

$$[L_{m\delta}, L_{n\delta}] = 2(m-n) K L_{(m+n)\delta}$$

$$[\mathcal{H}_\mu(x), \mathcal{H}_\nu(y)] \sim g(x) \partial\delta(x-y) \mathcal{H}_\lambda(y)$$



# Conclusions

- The “near cosmological singularity limit” suggests that there exists a “holographic” correspondence between  $D = 11$  SUGRA (or  $M$ -theory) and a  $D = 1$  spinning particle dynamics on the infinite-dimensional coset space  $E_{10}/K_{10}$ .
- The quantum dynamics of the coset model could provide both a background-independent formulation of  $M$ -theory, and a description of the (de-)emergence of space at a big-bang (big-crunch) singularity.
- There is an “arithmetic chaos” structure contained in the coset model (via Weyl group  $W(E_{10})$ ) that could play an important role.
- From the coset geodesic dynamics part, one expects the quantum wave function  $\Psi$  to be an automorphic function under  $E_{10}(\mathbb{Z})$ , solution of  $\square\Psi = 0$ .
- $\exists$  probably too many d.o.f. in the coset. One expects the existence of an infinite tower of constraints  $\mathcal{L}_\alpha \Psi = 0$ , possibly linked (for IIA) with the  $\Lambda_1$  representation and leading to a semi-open algebra of the type,  $[\mathcal{L}_\alpha, \mathcal{L}_\beta] = \sum_{\gamma} J_{\alpha+\beta-\gamma} \mathcal{L}_\gamma$ , which (hopefully) is a generalization of the diffeomorphism algebra:  $[\mathcal{H}_\mu(x), \mathcal{H}_\nu(y)] \sim g(x) \partial\delta(x-y) \mathcal{H}_\lambda(y)$ .