

# Asymptotic safety

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## A QUANTUM THEORY OF GRAVITY

Expand

$$g_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}$$

Gauge fix using a background gauge fixing condition e.g.

$$S_{GF} = \frac{1}{2} \int dx \sqrt{-\bar{g}} \bar{g}^{\mu\nu} \chi_\mu \chi_\nu ; \quad \chi_\mu = \bar{\nabla}^\nu h_{\nu\mu} - \frac{1}{2} \bar{\nabla}_\mu h$$

Add ghost Lagrangian

$$S_{ghost} = \int dx \sqrt{-\bar{g}} \bar{c}^\mu (-\delta_\mu^\nu \bar{\nabla}^2 - \bar{R}^\nu{}_\mu) c_\nu$$

Compute  $\Gamma(\bar{g}_{\mu\nu}, h_{\mu\nu})$ .

Formalism preserves background gauge invariance

$$\delta_\epsilon \bar{g}_{\mu\nu} = \bar{\nabla}_\mu \epsilon_\nu + \bar{\nabla}_\nu \epsilon_\mu, \quad \delta_\epsilon h_{\mu\nu} = \mathcal{L}_\epsilon h_{\mu\nu}.$$

## ISSUES

Non-renormalizable (Goroff and Sagnotti 1985)

- interaction strength grows like  $\tilde{G} = Gk^2$
- violation of unitarity
- lack of predictivity

## General EFT recipe

- fix the level of precision that is required in the calculation
- fix the ratio  $E/M$ . This determines the order of the expansion that will be required
- at the given order in the expansion, determine all the terms in the action that can contribute to the given process. By power counting there can only be finitely many.
- if the fundamental theory is known, their coefficients can in principle be calculated
- otherwise, measure them with as many experiments.
- use the EFT Lagrangian to compute the amplitudes to the desired precision

## Chiral perturbation theory

Strong interactions at low energy described by chiral model

$$S = \int dx \left[ \frac{f_\pi^2}{4} \text{tr}(U^{-1} \partial U)^2 + \ell_1 \text{tr}((U^{-1} \partial U)^2)^2 + \ell_2 \text{tr}((U^{-1} \partial U)^2)^2 + \mathcal{O}(\partial^6) \right]$$

Expansion parameter  $E/4\pi f_\pi$ .

For low energy meson physics need  $f_\pi$ ,  $\ell_1$ ,  $\ell_2$  and perhaps a few others. One loop in  $f_\pi$ , tree level in  $\ell_1$ ,  $\ell_2$ .

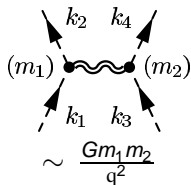
Successfully describe a rich phenomenology.

# Gravity

$$S = \int dx \sqrt{g} [2m_P^2 \Lambda - m_P^2 R + \ell_1 R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} + \ell_2 R_{\mu\nu} R^{\mu\nu} + \ell_3 R^2 + O(\partial^6)]$$

$m_P$  similar to  $f_\pi$ .

## Newtonian potential



$$V(r) = \int \frac{d^3 q}{(2\pi)^3} \frac{Gm_1 m_2}{q^2} e^{iqr} = -\frac{Gm_1 m_2}{r}$$

## Leading quantum correction

For dimensional reasons the leading correction is

$$V(r) = -\frac{Gm_1 m_2}{r} \left[ 1 + \alpha \frac{G(m_1 + m_2)}{rc^2} + \beta \frac{G\hbar}{r^2 c^3} + \dots \right]$$

and

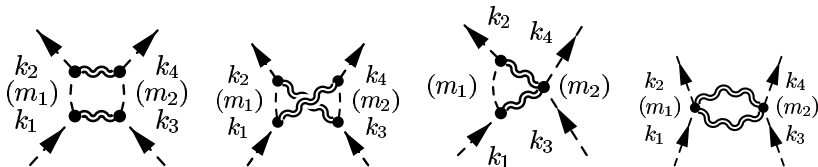
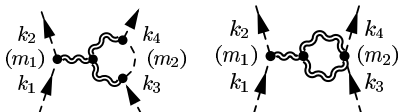
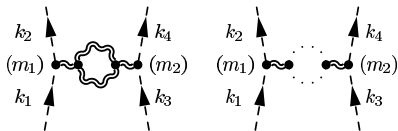
$$\int \frac{d^3 q}{(2\pi)^3} \log\left(\frac{q^2}{\mu^2}\right) e^{iqr} = -\frac{1}{2\pi^2 r^3}$$

Clearly distinct from contributions of local counterterms that give analytic corrections to the amplitude. E.g.

$$\int \frac{d^3 q}{(2\pi)^3} \ell e^{iqr} = \ell \delta(r)$$



# One loop graphs



## Evaluation

- J.F. Donoghue, P.R.L. 72, 2996 (1994); P.R.D50, 3874 (1994)
- H.W. Hamber, S. Liu, Phys. Lett. B357, 51 (1995)
- A. Akhundov, S. Bellucci, A. Shiekh, Phys. Lett. B395, 16 (1997)
- N.E.J. Bjerrum-Bohr (2002) Phys. Rev. D66, 084023
- I.B. Khriplovich, G.G. Kirilin (2002) J. Exp. Theor. Phys. 95, 981-986 (Zh. Eksp. Teor. Fiz. 95, 1139-1145 (2002))
- N.E.J. Bjerrum-Bohr, J.F. Donoghue, B.R. Holstein Phys. Rev. D68, 084005; Erratum-ibid.D71, 069904 (2005)
- N.E.J. Bjerrum-Bohr, J.F. Donoghue and B.R. Holstein Phys. Rev. D 67, 084033 (2003) [Erratum-ibid. D 71 (2005) 069903
- I.B. Khriplovich, G.G. Kirilin (2004) J. Exp. Theor. Phys. 98, 1063-1072

## A prediction of quantum gravity

$$V(r) = -\frac{Gm_1m_2}{r} \left[ 1 + 3\frac{G(m_1 + m_2)}{rc^2} + \frac{41}{10\pi} \frac{G\hbar}{r^2c^3} + \dots \right]$$

No renormalization ambiguities.

Probably the most reliable result in quantum gravity.

$M = M_\odot$	$\frac{GM_\odot}{rc^2}$	$\frac{G\hbar}{r^2c^3}$
$r = R_\odot$	$10^{-6}$	$10^{-88}$
$r = r_{S_\odot}$	0.5	$10^{-76}$

## Lessons

- no clash between QM and GR
- EFT is consistent and predictive at low enough energy
- experimentally indistinguishable from classical GR
- agrees with all experimental data
- to some extent, background independent
- open issues in the UV, IR, strong field

## Outline

- 1 **Philosophy**
- 2 **Asymptotic safety**
- 3 **1-loop gravity**
- 4 **FRGE**
- 5 **Conclusions**

## THE PROBLEMS OF QG, AND THE RG SOLUTION

Try to extend beyond Planck scale

- interaction strength grows like  $\tilde{G} = Gk^2$
- lack of predictivity

first problem solved if  $\tilde{G} = G(k)k^2 \rightarrow \tilde{G}_*$

more generally, if  $\Gamma_k(\bar{g}_{\mu\nu}, h_{\mu\nu}) = \sum_i g_i(k) \mathcal{O}_i(\bar{g}_{\mu\nu}, h_{\mu\nu})$

require  $\tilde{g}_i \rightarrow \tilde{g}_{i*}$ , where  $\tilde{g}_i = k^{-d_i} g_i$

## THE RG SOLUTION CONT'D

Define a theory space (fields, symmetry, action functionals) parameterized by  $\tilde{g}_j$ .

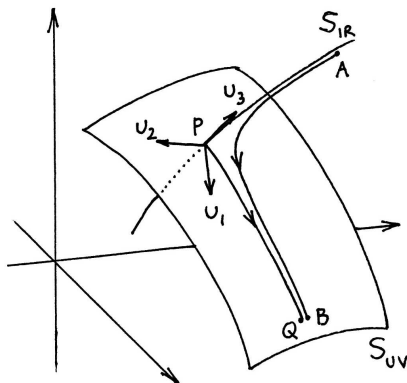
- RG trajectory is “renormalizable” or “Asymptotically Safe” if it flows to a FP in the UV
- The RG trajectories that flow into the FP for  $k \rightarrow \infty$  form the UV critical surface  $S_{UV}$
- Predictivity demands  $\dim(S_{UV})$  is finite

## EXAMPLE: QCD

- Gaussian Fixed Point at  $\tilde{g}_{i*} = 0$ .
- $\tilde{g}_i = k^{-d_i} g_i$
- $\tilde{\beta}_i = \partial_t \tilde{g}_i = -d_i \tilde{g}_i + k^{-d_i} \beta_i$
- $M_{ij}|_* = \frac{\partial \tilde{\beta}_i}{\partial \tilde{g}_j}|_* = -d_i \delta_{ij}$
- relevant couplings=renormalizable couplings



## GENERAL PICTURE



## ONE LOOP CORRECTIONS IN EINSTEIN'S THEORY

$$k \frac{d}{dk} \frac{1}{16\pi G(k)} = ck^{d-2}$$

$$k \frac{dG}{dk} = -16\pi c G^2 k^{d-2}$$

$$\tilde{G} = Gk^{d-2}$$

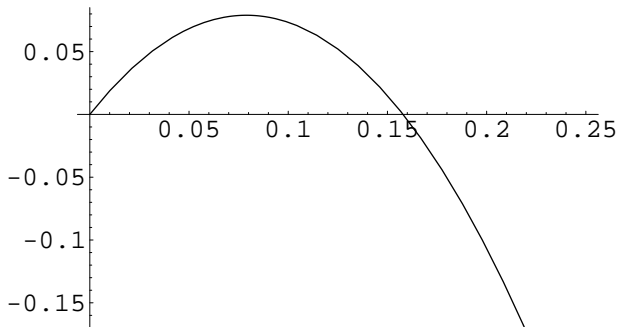
$$k \frac{d\tilde{G}}{dk} = (d-2)\tilde{G} - 16\pi c \tilde{G}^2$$

fixed point at  $\tilde{G} = (d-2)/16\pi c$

GRAVITY IN  $2 + \epsilon$ 

$$d - 2 = \epsilon$$

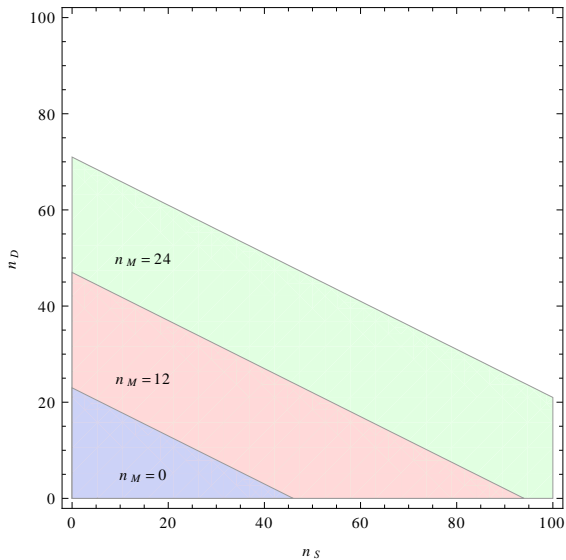
$$\tilde{G} = Gk^\epsilon$$
$$\beta_{\tilde{G}} = \epsilon \tilde{G} - \frac{38}{3} \tilde{G}^2$$



## ONE LOOP, $d = 4$

$$\mathbf{c} = \frac{11}{3\pi}, \frac{35}{8\pi}, \frac{23}{3\pi}, \dots$$

# ONE LOOP, WITH MATTER



## HIGHER DERIVATIVE GRAVITY

$$\Gamma_k = \int d^4x \sqrt{g} \left[ 2Z\Lambda - ZR + \frac{1}{2\lambda} \left( C^2 - \frac{2\omega}{3} R^2 + 2\theta E \right) \right]$$

$$Z = \frac{1}{16\pi G}$$

K.S. Stelle, Phys. Rev. **D16**, 953 (1977).

J. Julve, M. Tonin, Nuovo Cim. **46B**, 137 (1978).

E.S. Fradkin, A.A. Tseytlin, Phys. Lett. **104 B**, 377 (1981).

I.G. Avramidi, A.O. Barvinski, Phys. Lett. **159 B**, 269 (1985).

G. de Berredo–Peixoto and I. Shapiro, Phys.Rev. **D71** 064005 (2005).

A. Codello and R. P., Phys.Rev.Lett. **97** 22 (2006)

## BETA FUNCTIONS I

$$\beta_\lambda = -\frac{1}{(4\pi)^2} \frac{133}{10} \lambda^2$$

$$\beta_\omega = -\frac{1}{(4\pi)^2} \frac{25 + 1098\omega + 200\omega^2}{60} \lambda$$

$$\beta_\theta = \frac{1}{(4\pi)^2} \frac{7(56 - 171\theta)}{90} \lambda$$

$$\lambda(k) = \frac{\lambda_0}{1 + \lambda_0 \frac{1}{(4\pi)^2} \frac{133}{10} \log\left(\frac{k}{k_0}\right)}$$

$$\omega(k) \rightarrow \omega_* \approx -0.0228$$

$$\theta(k) \rightarrow \theta_* \approx 0.327$$

## BETA FUNCTIONS II

$$\beta_{\tilde{\Lambda}} = -2\tilde{\Lambda} + \frac{1}{(4\pi)^2} \left[ \frac{1 + 20\omega^2}{256\pi\tilde{G}\omega^2} \lambda^2 + \frac{1 + 86\omega + 40\omega^2}{12\omega} \lambda\tilde{\Lambda} \right]$$

$$- \frac{1 + 10\omega^2}{64\pi^2\omega} \lambda + \frac{2\tilde{G}}{\pi} - q(\omega)\tilde{G}\tilde{\Lambda}$$

$$\beta_{\tilde{G}} = 2\tilde{G} - \frac{1}{(4\pi)^2} \frac{3 + 26\omega - 40\omega^2}{12\omega} \lambda\tilde{G} - q(\omega)\tilde{G}^2$$

where  $q(\omega) = (83 + 70\omega + 8\omega^2)/18\pi$

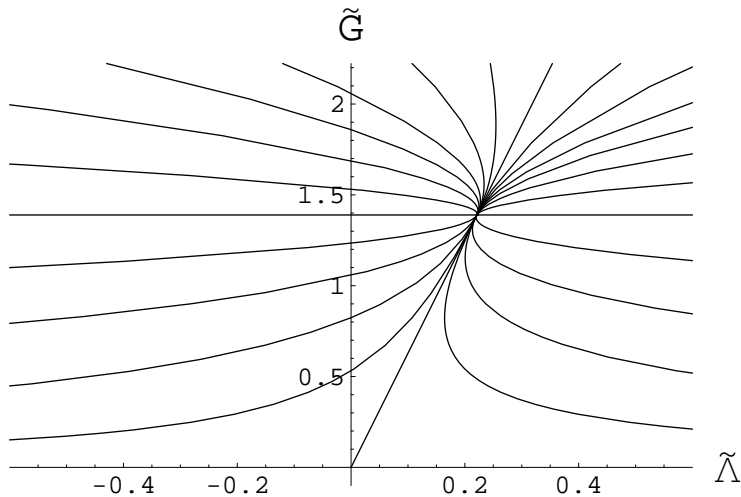


FLOW IN  $\tilde{\Lambda}$ - $\tilde{G}$  PLANE I

$$\beta_{\tilde{\Lambda}} = -2\tilde{\Lambda} + \frac{2\tilde{G}}{\pi} - q_* \tilde{G} \tilde{\Lambda}$$
$$\beta_{\tilde{G}} = 2\tilde{G} - q_* \tilde{G}^2$$

where  $q_* = q(\omega_*) \approx 1.440$

$$\tilde{\Lambda}_* = \frac{1}{\pi q_*} \approx 0.221, \quad \tilde{G}_* = \frac{2}{q_*} \approx 1.389.$$

FLOW IN  $\tilde{\Lambda}$ - $\tilde{G}$  PLANE II

## Topologically massive gravity

### Action

$$S(g) = Z \int d^3x \sqrt{g} \left( -2\Lambda + R + \frac{1}{2\mu} \varepsilon^{\lambda\mu\nu} \Gamma_{\lambda\sigma}^{\rho} \left( \partial_{\mu} \Gamma_{\nu\rho}^{\sigma} + \frac{2}{3} \Gamma_{\mu\tau}^{\sigma} \Gamma_{\nu\rho}^{\tau} \right) \right)$$

$$Z = \frac{1}{16\pi G}$$

### Dimensionless combinations of couplings

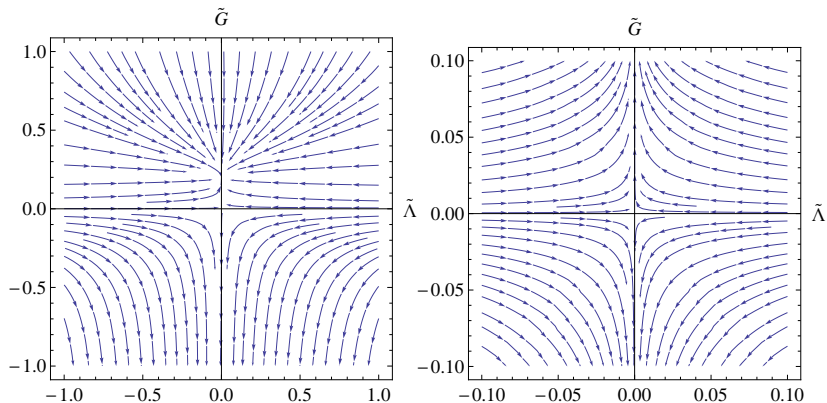
$$\nu = \mu G ; \quad \tau = \Lambda G^2 ; \quad \phi = \mu / \sqrt{|\Lambda|}$$

## Beta functions of

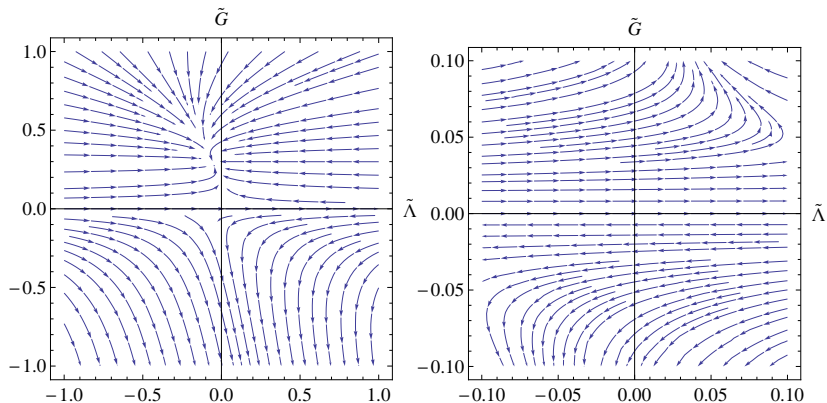
$$\begin{aligned}\beta_\nu &= 0, \\ \beta_{\tilde{G}} &= \tilde{G} + B(\tilde{\mu})\tilde{G}^2, \\ \beta_{\tilde{\Lambda}} &= -2\tilde{\Lambda} + \frac{1}{2}\tilde{G} \left( A(\tilde{\mu}, \tilde{\Lambda}) + 2B(\tilde{\mu})\tilde{\Lambda} \right)\end{aligned}$$

**Since  $\nu = \mu G = \tilde{\mu}\tilde{G}$  is constant**

can replace  $\tilde{\mu}$  by  $\nu/\tilde{G}$



**Figure :** The flow in the  $\tilde{\Lambda}$ - $\tilde{G}$  plane for  $\alpha = 0$ ,  $\nu = 5$ . Right: enlargement of the region around the origin, showing the Gaussian FP. The beta functions become singular at  $|\tilde{G}| = 1.9245$ .



**Figure :** The flow in the  $\tilde{\Lambda}$ - $\tilde{G}$  plane for  $\alpha = 0$ ,  $\nu = 0.1$ . Right: enlargement of the region around the origin, showing that there is no Gaussian FP. The beta functions diverge on the  $\tilde{\Lambda}$  axis.

Recently extended to TM SUGRA

[R.P., C. Pope, M. Perry, E. Sezgin,  
arXiv 1302.0868]

## FRGE

Define  $\Gamma_k(\bar{g}_{\mu\nu}, h_{\mu\nu})$ . It satisfies a simple functional differential equation

$$k \frac{d\Gamma_k}{dk} = \beta(\bar{g}_{\mu\nu}, h_{\mu\nu})$$

The quantity  $\beta$  is UV and IR finite.

Since  $\lim_{k \rightarrow 0} \Gamma_k = \Gamma$ , can use FRGE to calculate the effective action.

Single-field truncations:

$$\Gamma_k(g_{\mu\nu}) = \Gamma_k(g_{\mu\nu}, 0)$$



## EINSTEIN-HILBERT TRUNCATION I

$$\Gamma_k(g) = \frac{1}{16\pi G_k} \int dx \sqrt{g} (2\Lambda_k - R) + S_{GF} + S_{ghost}$$

with  $h_{\mu\nu} \rightarrow Z_h^{1/2} h_{\mu\nu}$ ,  $c^\mu \rightarrow Z_c^{1/2} c^\mu$  and  $c_\nu \rightarrow Z_c^{1/2} c_\nu$

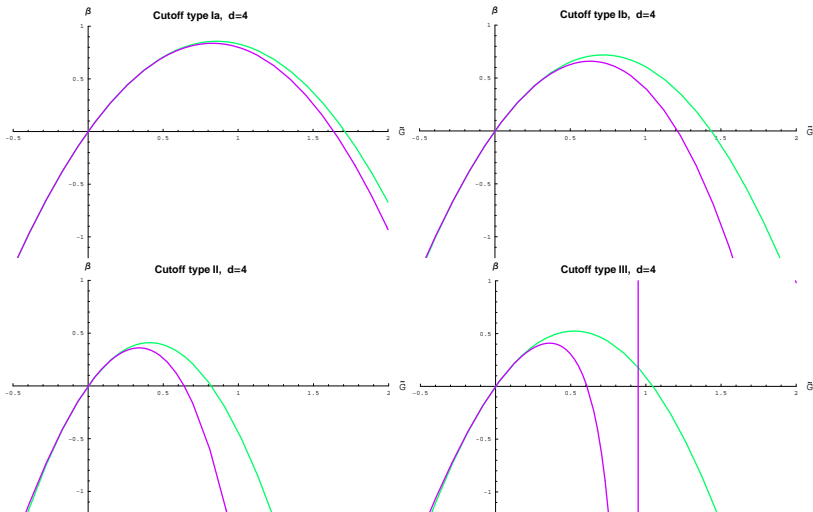
where  $\eta_h = -\frac{\partial_t Z_h}{Z_h}$ ,  $\eta_c = -\frac{\partial_t Z_c}{Z_c}$

Assume  $\eta_h = -\frac{\partial_t Z}{Z}$ ,  $\eta_c = 0$ .

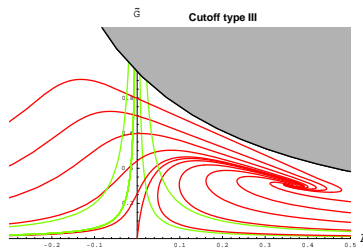
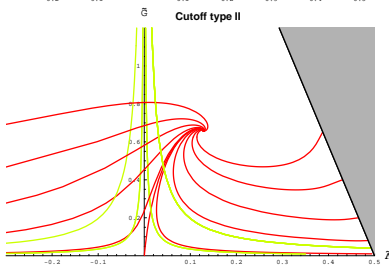
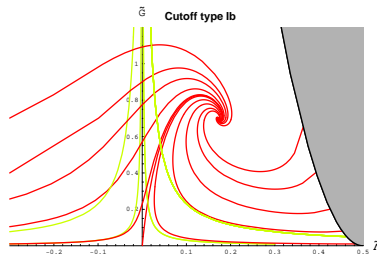
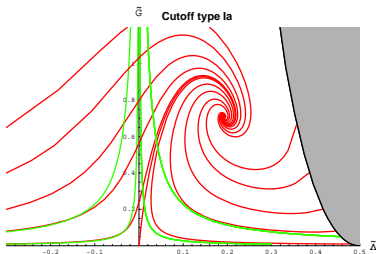
$$\beta_{\tilde{\Lambda}} = \frac{-2(1 - 2\tilde{\Lambda})^2 \tilde{\Lambda} + \frac{36 - 41\tilde{\Lambda} + 42\tilde{\Lambda}^2 - 600\tilde{\Lambda}^3}{72\pi} \tilde{G} + \frac{467 - 572\tilde{\Lambda}}{288\pi^2} \tilde{G}^2}{(1 - 2\tilde{\Lambda})^2 - \frac{29 - 9\tilde{\Lambda}}{72\pi} \tilde{G}}$$

$$\beta_{\tilde{G}} = \frac{2(1 - 2\tilde{\Lambda})^2 \tilde{G} - \frac{373 - 654\tilde{\Lambda} + 600\tilde{\Lambda}^2}{72\pi} \tilde{G}^2}{(1 - 2\tilde{\Lambda})^2 - \frac{29 - 9\tilde{\Lambda}}{72\pi} \tilde{G}}$$

# EINSTEIN-HILBERT TRUNCATION II



## EINSTEIN–HILBERT TRUNCATION III



## FOURTH ORDER GRAVITY

- $R^2$  O.Lauscher, M. Reuter, Phys. Rev. D 66, 025026 (2002) arXiv:hep-th/0205062
- $R^2 + C^2$  D. Benedetti, P.F. Machado, F. Saueressig, Mod. Phys. Lett. A24, 2233-2241 (2009) arXiv:0901.2984 [hep-th] Nucl. Phys. B824, 168-191 (2010), arXiv:0902.4630 [hep-th]
- M. Niedermaier, Nucl. Phys. B833, 226-270 (2010)

$f(R)$ 

talk by Dario Benedetti

## BI-FIELD TRUNCATION I

$$\Gamma_k(\bar{g}_{\mu\nu}, h_{\mu\nu}) = \frac{1}{16\pi G_k} \int dx \sqrt{g} (2\Lambda_k - R) + S_{GF} + S_{ghost}$$

flow  $G_k, \Lambda_k, Z_h, Z_c$ .

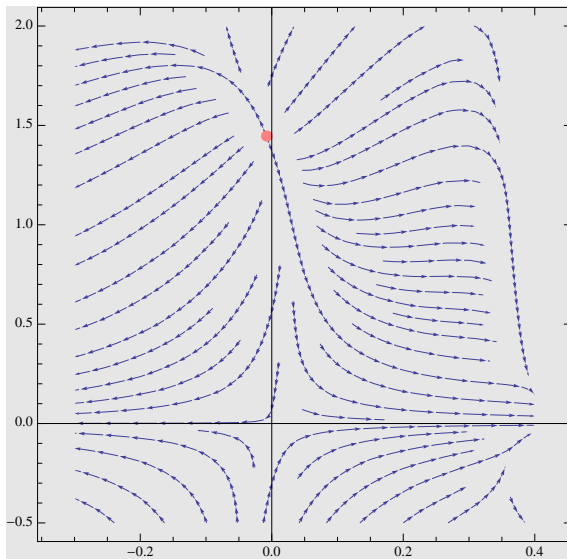
FP at  $\tilde{\Lambda} = -0.00812786, \tilde{G} = 1.44627$

scaling exponents  $-3.32286, -1.95403$

anomalous dimensions  $\eta_h = 0.0720708, \eta_c = -1.50325$ .

[A. Codello, G. d'Odorico, G. Pagani, in preparation]

## BI-FIELD TRUNCATION II



## SUMMARY

- AS of gravity by now quite plausible
- use continuum covariant QFT
- bottom up approach, use EFT framework
- almost guaranteed to give correct low energy limit
- hints of agreement with CDT



## OUTLOOK

- more complicated/complete truncations
- matter coupled to gravity
- observables
- describe as CFT